

1. Integrating a Scalar Along a Parabola. Let C be the portion of the parabola $y = x^2$ between points $(0, 0)$ and $(1, 1)$, and consider the scalar function $f(x, y) = x$. Compute the integral of f along C :

$$\int_C f \, ds.$$

[Hint: You must choose a parametrization of C . I recommend $\mathbf{r}(t) = (t, t^2)$ with $0 \leq t \leq 1$. The resulting integral may be computed by hand using substitution.]

2. Projection. Let \mathbf{F} and \mathbf{u} be any vectors with $\|\mathbf{u}\| = 1$ and let \mathbf{p} be the *component of \mathbf{F} in the direction of the unit vector \mathbf{u}* .

- Since \mathbf{p} is parallel to \mathbf{u} we know that $\mathbf{p} = t\mathbf{u}$ for some scalar t . Use the fact that the vector $\mathbf{p} - \mathbf{F}$ is perpendicular to \mathbf{u} to prove that $t = \mathbf{F} \bullet \mathbf{u}$.
- Draw a picture showing the vectors \mathbf{F} , \mathbf{u} and \mathbf{p} .

3. Integrating Vector Fields Along Different Curves. By definition, the integral of a vector field \mathbf{F} along a parametrized curve $\mathbf{r}(t)$ is

$$\int \mathbf{F}(\mathbf{r}(t)) \bullet \mathbf{r}'(t) \, dt.$$

Consider the two fields $\mathbf{F}(x, y) = \langle 2xy, x^2 \rangle$, $\mathbf{G}(x, y) = \langle 2y, x^2 \rangle$, and the two different curves $\mathbf{r}(t) = (t, t)$ and $\mathbf{s}(t) = (t, t^2)$ between the points $(0, 0)$ and $(1, 1)$.

- Integrate \mathbf{F} along $\mathbf{r}(t)$ and $\mathbf{s}(t)$. Observe that you get the same answer.
- Integrate \mathbf{G} along $\mathbf{r}(t)$ and $\mathbf{s}(t)$. Observe that you don't get the same answer.

4. Area of a Pringle. Let D be the surface in \mathbb{R}^3 defined by $z = xy$ and $x^2 + y^2 \leq 1$, which looks like a pringle chip. We can parametrize this region by

$$\mathbf{r}(u, v) = \langle u \cos v, u \sin v, (u \cos v)(u \sin v) \rangle$$

with $0 \leq u \leq 1$ and $0 \leq v \leq 2\pi$.

- Compute the tangent vectors \mathbf{r}_u and \mathbf{r}_v . [Hint: Use the identity $\sin(2v) = 2 \sin v \cos v$.]
- Compute the cross product $\mathbf{r}_u \times \mathbf{r}_v$.
- Compute the length $\|\mathbf{r}_u \times \mathbf{r}_v\|$ and simplify as much as possible. [Hint: The answer is $\|\mathbf{r}_u \times \mathbf{r}_v\| = u\sqrt{u^2 + 1}$.]
- Use your answer from part (c) to compute the area of the pringle:

$$\text{Area}(D) = \iint_D 1 \|\mathbf{r}_u \times \mathbf{r}_v\| \, dudv.$$

5. Proof of Conservation of Energy. A *conservative force field* \mathbf{F} has the form $\mathbf{F} = -\nabla f$ for some *scalar potential* f . Suppose that a particle of mass m travels along a trajectory $\mathbf{r}(t)$. Newton says that the force \mathbf{F} acting on the particle satisfies

$$\mathbf{F}(\mathbf{r}(t)) = m\mathbf{r}''(t).$$

The *kinetic energy* of the particle at time t is $\text{KE}(t) = \frac{1}{2}m\|\mathbf{r}'(t)\|^2$, the *potential energy* at time t is $\text{PE}(t) = f(\mathbf{r}(t))$, and the *total mechanical energy* is $E(t) = \text{KE}(t) + \text{PE}(t)$. Use the chain rule and product rule for derivatives to show that

$$E'(t) = 0.$$

[Hint: Write $\|\mathbf{r}'(t)\|^2 = \mathbf{r}'(t) \bullet \mathbf{r}'(t)$.]

6. Application of Conservation of Energy. Choose a coordinate system in \mathbb{R}^3 with the sun at position $(0, 0, 0)$. Suppose that the sun has mass M . If $\mathbf{F}(x, y, z)$ is the gravitational force exerted by the sun on a spaceship of mass m at position (x, y, z) , Newton tells us that $\mathbf{F}(x, y, z) = -\nabla f(x, y, z)$, where¹

$$f(x, y, z) = -1 \cdot \frac{GMm}{\sqrt{x^2 + y^2 + z^2}}$$

is called the *gravitational potential*. At a certain time, the spaceship has speed s_0 and distance d_0 from the origin. At a later time the spaceship has speed s_1 and distance d_1 from the origin. Use conservation of energy to compute s_1 in terms of s_0 , d_0 and d_1 . (Assume that no other forces are acting on the spaceship.)

¹ G is a constant of nature called the *gravitational constant*.