

1. Various Kinds of First and Second Derivatives in \mathbb{R}^3 . For any scalar field $f(x, y, z)$ we define a vector field $\text{grad}(f)$ and a scalar field laplacian(f) by

$$\begin{aligned}\text{grad}(f) &= \text{“}\nabla f\text{”} = \langle f_x, f_y, f_z \rangle, \\ \text{laplacian}(f) &= \text{“}\nabla^2 f\text{”} = f_{xx} + f_{yy} + f_{zz}.\end{aligned}$$

and for any vector field $\mathbf{F}(x, y, z) = \langle P(x, y, z), Q(x, y, z), R(x, y, z) \rangle$ we define a vector field $\text{curl}(\mathbf{F})$ and a scalar field $\text{div}(\mathbf{F})$ by

$$\begin{aligned}\text{curl}(\mathbf{F}) &= \text{“}\nabla \times \mathbf{F}\text{”} = \langle R_y - Q_z, P_z - R_x, Q_x - P_y \rangle, \\ \text{div}(\mathbf{F}) &= \text{“}\nabla \bullet \mathbf{F}\text{”} = P_x + Q_y + R_z.\end{aligned}$$

- (a) For any scalar field $f : \mathbb{R}^3 \rightarrow \mathbb{R}$ check that $\text{curl}(\text{grad}(f)) = \langle 0, 0, 0 \rangle$.
- (b) For any vector field $\mathbf{F} : \mathbb{R}^3 \rightarrow \mathbb{R}^3$ check that $\text{div}(\text{curl}(\mathbf{F})) = 0$.
- (c) For any scalar field $f : \mathbb{R}^3 \rightarrow \mathbb{R}$ check that $\text{div}(\text{grad}(f)) = \text{laplacian}(f)$.

(a): Consider any scalar field f and its gradient vector field $\text{grad}(f) = \langle f_x, f_y, f_z \rangle$. In order to compute the curl of $\text{grad}(f)$ we will write $\text{grad}(f) = \langle P, Q, R \rangle$, so that $P = f_x$, $Q = f_y$ and $R = f_z$. Then by definition of the curl we have

$$\begin{aligned}\text{curl}(\text{grad}(f)) &= \langle R_y - Q_z, P_z - R_x, Q_x - P_y \rangle, \\ &= \langle (f_z)_y - (f_y)_z, (f_x)_z - (f_z)_x, (f_y)_x - (f_x)_y \rangle, \\ &= \langle f_{zy} - f_{yz}, f_{xz} - f_{zx}, f_{yx} - f_{xy} \rangle,\end{aligned}$$

which simplifies to $\langle 0, 0, 0 \rangle$ because the mixed partial derivatives of f commute.

(b): Consider any vector field $\mathbf{F} = \langle P, Q, R \rangle$ and its curl

$$\text{curl}(\mathbf{F}) = \langle R_y - Q_z, P_z - R_x, Q_x - P_y \rangle.$$

Then by definition the divergence of $\text{curl}(\mathbf{F})$ is

$$\begin{aligned}\text{div}(\text{curl}(\mathbf{F})) &= (R_y - Q_z)_x + (P_z - R_x)_y + (Q_x - P_y)_z \\ &= (R_y)_x - (Q_z)_x + (P_z)_y - (R_x)_y + (Q_x)_z - (P_y)_z \\ &= R_{yx} - Q_{zx} + P_{zy} - R_{xy} + Q_{xz} - P_{yz} \\ &= (P_{zy} - P_{yz}) + (Q_{xz} - Q_{zx}) + (R_{yx} - R_{xy}),\end{aligned}$$

which simplifies to $0 + 0 + 0 = 0$ because the mixed partial derivatives of P, Q, R commute.

Remark: These are the three most basic “derivative identities” of vector calculus. They are important for such topics as fluid dynamics and electro-magnetism.

(c): Consider any scalar field f and its gradient vector field $\text{grad}(f) = \langle f_x, f_y, f_z \rangle$. Then from the definition of divergence and laplacian we have

$$\text{div}(\text{grad}(f)) = (f_x)_x + (f_y)_y + (f_z)_z = f_{xx} + f_{yy} + f_{zz} = \text{laplacian}(f).$$

2. Conservative Vector Fields. Consider the vector field

$$\mathbf{F}(x, y, z) = \langle y + z, x + z, x + y \rangle.$$

- (a) Check that the curl is constantly zero: $\nabla \times \mathbf{F}(x, y, z) = \langle 0, 0, 0 \rangle$.

- (b) It follows from part (a) that there exists some scalar field $f(x, y, z)$ satisfying $\nabla f = \mathbf{F}$. Find one such scalar field. [Hint: Integrate \mathbf{F} along an arbitrary path starting at some arbitrary point and ending at (x, y, z) .]

(a): Let $\mathbf{F} = \langle y + z, x + z, x + y \rangle = \langle P, Q, R \rangle$. Then we have

$$\begin{aligned}\nabla \times \mathbf{F} &= \langle R_y - Q_z, P_z - R_x, Q_x - P_y \rangle \\ &= \langle (x + y)_y - (x + z)_z, (y + z)_z - (x + y)_x, (x + z)_x - (y + z)_y \rangle \\ &= \langle 1 - 1, 1 - 1, 1 - 1 \rangle \\ &= \langle 0, 0, 0 \rangle.\end{aligned}$$

Remark: If $\mathbf{F} = \nabla f$ for some scalar field f then we know from Problem 1(a) that $\nabla \times \mathbf{F} = \langle 0, 0, 0 \rangle$. The “fundamental theorem of conservative vector fields” says that the converse is also true. That is, if $\nabla \times \mathbf{F} = \langle 0, 0, 0 \rangle$ (and if \mathbf{F} is continuous everywhere in \mathbb{R}^3 , as it is in this example) then there must exist some scalar field f such that $\mathbf{F} = \nabla f$.

(b): In order to find such a scalar field we can just take $f = \int \mathbf{F} \bullet \mathbf{T} ds$, where the line integral is computed over **any path with endpoint** (x, y, z) .¹ The easiest imaginable such path is $\mathbf{r}(t) = \langle xt, yt, zt \rangle$ for $0 \leq t \leq 1$. Thus we define

$$\begin{aligned}f(x, y, z) &= \int_0^1 \mathbf{F}(xt, yt, zt) \bullet \langle xt, yt, zt \rangle' dt \\ &= \int_0^1 \langle yt + zt, xt + zt, xt + yt \rangle \bullet \langle x, y, z \rangle dt \\ &= \int_0^1 [(yt + zt)x + (xt + zt)y + (xt + yt)z] dt \\ &= \int_0^1 [xyt + xzt + xyt + zyt + xzt + yzt] dt \\ &= 2(xy + xz + yz) \cdot \int_0^1 t dt \\ &= 2(xy + xz + yz) \cdot \left[\frac{1^2}{2} - \frac{0^2}{2} \right] \\ &= xy + xz + yz.\end{aligned}$$

We verify that this satisfies the desired property $\nabla f = \langle y + z, x + z, x + y \rangle$:

$$\begin{aligned}\nabla(xy + xz + yz) &= \langle (xy + xz + yz)_x, (xy + xz + yz)_y, (xy + xz + yz)_z \rangle \\ &= \langle y + z + 0, x + 0 + z, 0 + x + y \rangle \\ &= \langle y + z, x + z, x + y \rangle.\end{aligned}$$

3. Gravitational Potential. A sun of mass M sits at the origin in \mathbb{R}^3 . According to Newton, the gravitational force due to the sun acting on a particle of mass m at the point (x, y, z) has the form

$$\mathbf{F}(x, y, z) = \frac{-GMm}{(x^2 + y^2 + z^2)^{3/2}} \cdot \langle x, y, z \rangle,$$

where G is the gravitational constant.

¹This is the three-dimensional version of the Calculus I theorem that $g(x) = \int_a^x G(t) dt$ satisfies $g'(x) = G(x)$ for **any lower endpoint** a .

(a) Check that the following scalar field $f : \mathbb{R}^3 \rightarrow \mathbb{R}$ satisfies $\nabla f = \mathbf{F}$:

$$f(x, y, z) = \frac{+GMm}{\sqrt{x^2 + y^2 + z^2}}.$$

(b) We can think of $U = -f$ as the *gravitational potential energy*. Suppose that our particle is held at rest at a position with distance D from the origin. At time zero the particle is allowed to fall towards the sun. If the sun has radius R , use conservation of energy to compute the particle's speed when it hits the sun's surface. Your answer will involve the constants G, M, R and D (but not m). [Assume that $D > R$.]

(a): First we compute the partial derivative of f with respect to x :

$$\begin{aligned} f_x &= \frac{\partial}{\partial x} GMm(x^2 + y^2 + z^2)^{-1/2} \\ &= GMm \cdot \frac{-1}{2} (x^2 + y^2 + z^2)^{-3/2} (2x + 0 + 0) \\ &= -GMmx / (x^2 + y^2 + z^2)^{3/2}. \end{aligned}$$

Identical computations show that $f_y = -GMmy / (x^2 + y^2 + z^2)^{3/2}$ and $f_z = -GMmz / (x^2 + y^2 + z^2)^{3/2}$, hence

$$\begin{aligned} \nabla f &= \langle f_x, f_y, f_z \rangle \\ &= \left\langle \frac{-GMmx}{(x^2 + y^2 + z^2)^{3/2}}, \frac{-GMmy}{(x^2 + y^2 + z^2)^{3/2}}, \frac{-GMmz}{(x^2 + y^2 + z^2)^{3/2}} \right\rangle \\ &= \frac{-GMm}{(x^2 + y^2 + z^2)^{3/2}} \cdot \langle x, y, z \rangle, \end{aligned}$$

as desired.

(b): If a particle with mass m and trajectory $\mathbf{r}(t)$ moves through the force field $\mathbf{F} = \nabla f$ then the theorem on conservation of energy says that the total mechanical energy is constant:

$$\text{KE} + \text{PE} = \frac{1}{2}m\|\mathbf{r}'(t)\|^2 - f(\mathbf{r}(t)) = \text{constant}.$$

Note that $f(\mathbf{r}(t)) = GMm / \sqrt{x(t)^2 + y(t)^2 + z(t)^2} = GMm / \|\mathbf{r}(t)\|$, where $\|\mathbf{r}(t)\|$ is just the distance from the position $\mathbf{r}(t)$ to the origin. Since the particle starts at rest with distance D to the origin we have

$$(\text{KE} + \text{PE})(\text{start}) = 0 - GMm/D.$$

Let v be the particle's speed when it hits the surface of the sun, i.e., when $\|\mathbf{r}(t)\| = R$. At this moment we have

$$(\text{KE} + \text{PE})(\text{hit}) = \frac{1}{2}mv^2 - GMm/R.$$

Finally, conservation of energy says that

$$\begin{aligned} (\text{KE} + \text{PE})(\text{hit}) &= (\text{KE} + \text{PE})(\text{start}) \\ \frac{1}{2}mv^2 - GMm/R &= 0 - GMm/D \\ v^2 &= 2GM/R - 2GM/D \\ v &= \sqrt{\frac{2GM}{R} - \frac{2GM}{D}}. \end{aligned}$$

Note that the mass m dropped out of the equation.

Remark: This equation works well for NASA, but it breaks down in extreme cases. For example, it predicts that $v \rightarrow \infty$ for a sun with large mass M and tiny radius R .

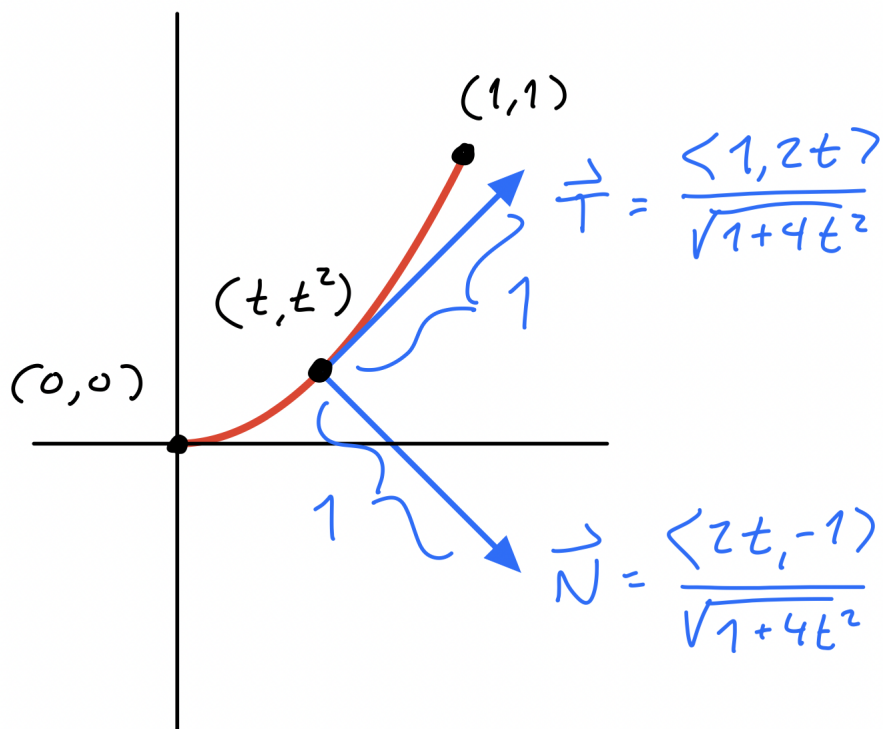
4. Line Integrals and Flux Integrals in \mathbb{R}^2 . Let C be the parametrized path $\mathbf{r}(t) = (t, t^2)$ for $0 \leq t \leq 1$, with velocity vector $\mathbf{r}'(t) = \langle 1, 2t \rangle$. From this parametrization we can define a unit tangent vector and a unit normal vector to the curve C at the point $\mathbf{r}(t)$:

$$\mathbf{T} = \frac{\langle 1, 2t \rangle}{\sqrt{1 + 4t^2}} \quad \text{and} \quad \mathbf{N} = \frac{\langle 2t, -1 \rangle}{\sqrt{1 + 4t^2}}.$$

Now consider the constant vector field $\mathbf{F}(x, y) = \langle 3, 1 \rangle$.

- Compute the line integral $\int_C \mathbf{F} \cdot \mathbf{T} ds$.
- Compute the flux integral $\int_C \mathbf{F} \cdot \mathbf{N} ds$.

Here is a picture of the path $\mathbf{r}(t) = (t, t^2)$ for $0 \leq t \leq 1$ together with the unit tangent and normal vectors at the point (t, t^2) :



Note that $\mathbf{T} = \mathbf{r}'(t)/\|\mathbf{r}'(t)\|$ is just velocity vector $\mathbf{r}'(t) = \langle 1, 2t \rangle$ divided by its magnitude $\|\mathbf{r}'(t)\| = \sqrt{1^2 + (2t)^2}$. The normal vector \mathbf{N} is just \mathbf{T} rotated 90° clockwise. We do this so that \mathbf{N} points “to the right” of the curve. When $\mathbf{r}(t)$ is the oriented boundary curve of a 2D region D this ensures that \mathbf{N} points “out of” the region. (The region D is always “to the left” of its boundary curve ∂D .)

(a): The line integral is

$$\begin{aligned} \int_C \mathbf{F} \cdot \mathbf{T} ds &= \int_0^1 \mathbf{F}(t, t^2) \cdot \frac{\langle 1, 2t \rangle}{\|\mathbf{r}'(t)\|} \|\mathbf{r}'(t)\| dt \\ &= \int_0^1 \mathbf{F}(t, t^2) \cdot \langle 1, 2t \rangle dt \end{aligned}$$

$$\begin{aligned}
&= \int_0^1 \langle 3, 1 \rangle \bullet \langle 1, 2t \rangle dt \\
&= \int_0^1 (3 + 2t) dt \\
&= \left[3t + 2\frac{t^2}{2} \right]_0^1 \\
&= 3 + 1 \\
&= 4.
\end{aligned}$$

(a): The flux integral is

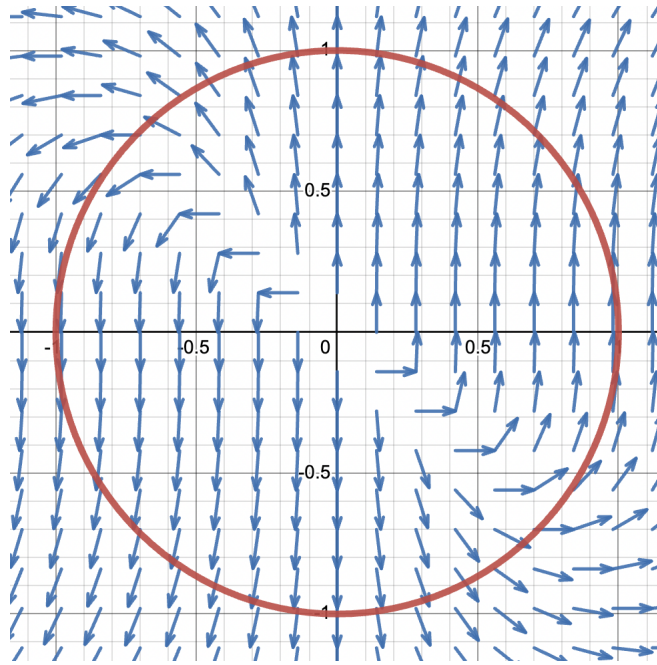
$$\begin{aligned}
\int \mathbf{F} \bullet \mathbf{N} ds &= \int_0^1 \mathbf{F}(t, t^2) \bullet \frac{\langle 2t, -1 \rangle}{\|\mathbf{r}'(t)\|} \|\mathbf{r}'(t)\| dt \\
&= \int_0^1 \mathbf{F}(t, t^2) \bullet \langle 2t, -1 \rangle dt \\
&= \int_0^1 \langle 3, 1 \rangle \bullet \langle 2t, -1 \rangle dt \\
&= \int_0^1 (6t - 1) dt \\
&= \left[6\frac{t^2}{2} - t \right]_0^1 \\
&= 3 - 1 \\
&= 2.
\end{aligned}$$

Remark: If we are flying a plane along the trajectory $\mathbf{r}(t)$ in a constant wind $\langle 3, 1 \rangle$, then the wind contributes 4 units of energy to our forward motion and $5/2$ units of energy trying to push us to the right.

5. Green's Theorem on a Circle. Let D be the unit disk in \mathbb{R}^2 centered at $(0, 0)$. Consider the vector field $\mathbf{F} = \langle P, Q \rangle = \langle xy^2, x + y \rangle$ with $\text{curl}(\mathbf{F}) = Q_x - P_y = 1 - 2xy$.

- Compute the integral $\iint_D \text{curl}(\mathbf{F}) dA$. [Hint: Polar coordinates are easiest. You may use the trigonometric identity $\sin(2\theta) = 2 \sin \theta \cos \theta$.]
- The boundary curve ∂D is the unit circle, oriented counterclockwise. Use the standard parametrization $\mathbf{r}(t) = (\cos t, \sin t)$ with $0 \leq t \leq 2\pi$ to set up the integral $\oint_{\partial D} \mathbf{F} \bullet \mathbf{T} ds$. You will probably not be able to evaluate this integral by hand. Use a computer to verify that you get the same answer as in part (a).

Here is a picture of the vector field $\mathbf{F} = \langle xy^2, x + y \rangle$ and the unit circle:



(a): First we integrate the scalar² $\text{curl}(\mathbf{F}) = Q_x - P_y = 1 - 2xy$ over the interior of the circle. We use polar coordinates $x = r \cos \theta$ and $y = r \sin \theta$ with $dA = dx dy = r dr d\theta$ to get

$$\begin{aligned}
 \iint_D \text{curl}(\mathbf{F}) dA &= \iint_D (1 + 2xy) dx dy \\
 &= \iint_D (1 + 2r \cos \theta r \sin \theta) r dr d\theta \\
 &= \iint_D [r + r^2 \sin(2\theta)] dr d\theta && 2 \cos \theta \sin \theta = \sin(2\theta) \\
 &= \int_0^{2\pi} \left(\int_0^1 [r + r^2 \sin(2\theta)] dr \right) d\theta \\
 &= \int_0^{2\pi} \left[\frac{r^2}{2} + \frac{r^3}{3} \cdot \sin(2\theta) \right]_0^1 d\theta \\
 &= \int_{\theta=0}^{\theta=2\pi} \left[\frac{1}{2} + \frac{1}{3} \sin(2\theta) \right] d\theta \\
 &= \int_{u=0}^{u=4\pi} \left[\frac{1}{2} + \frac{1}{3} \sin u \right] \frac{du}{2} && u = 2\theta, du = 2d\theta \\
 &= \frac{1}{2} \cdot \left[\frac{1}{2} \cdot u - \frac{1}{3} \cos u \right]_0^{4\pi} \\
 &= \frac{1}{2} \cdot \left[\frac{1}{2} \cdot 4\pi - 0 - \frac{1}{3} + \frac{1}{3} \right] \\
 &= \pi.
 \end{aligned}$$

²The curl in 3D is a vector field. The curl in 2D is just a scalar field.

(b): Now we integrate the field $\mathbf{F} = \langle xy^2, x + y \rangle$ along the boundary curve $\mathbf{r}(t) = \langle \cos t, \sin t \rangle$ for $0 \leq t \leq 2\pi$. Stokes Theorem says that we must get π as in part (a):

$$\begin{aligned}
 \int_{\partial D} \mathbf{F} \bullet \mathbf{T} \, ds &= \int_0^{2\pi} \mathbf{F}(\mathbf{r}(t)) \bullet \frac{\mathbf{r}'(t)}{\|\mathbf{r}'(t)\|} \cdot \|\mathbf{r}'(t)\| \, dt \\
 &= \int_0^{2\pi} \mathbf{F}(\mathbf{r}(t)) \bullet \mathbf{r}'(t) \, dt \\
 &= \int_0^{2\pi} \mathbf{F}(\cos t, \sin t) \bullet \langle -\sin t, \cos t \rangle \, dt \\
 &= \int_0^{2\pi} \langle \cos t \sin^2 t, \cos t + \sin t \rangle \bullet \langle -\sin t, \cos t \rangle \, dt \\
 &= \int_0^{2\pi} [(\cos t \sin^2 t)(-\sin t) + (\cos t + \sin t)(\cos t)] \, dt \\
 &= \int_0^{2\pi} [-\cos t \sin^2 t + \cos^2 t + \cos t \sin t] \, dt \\
 &\quad \vdots \\
 &= \pi.
 \end{aligned}$$

This can be solved using integration by parts and various trig identities, but I used a computer.

6. Stokes' Theorem on a Parabolic Dome. Let D be the two-dimensional surface in \mathbb{R}^3 defined by $z = 1 - x^2 - y^2$ and $z \geq 0$. This surface can be parametrized by

$$\mathbf{r}(u, v) = \langle u \cos v, u \sin v, 1 - u^2 \rangle \quad \text{with } 0 \leq u \leq 1 \text{ and } 0 \leq v \leq 2\pi.$$

The boundary curve ∂D is the unit circle in the x, y -plane, oriented counterclockwise, which can be parametrized as $\mathbf{r}(t) = \langle \cos t, \sin t, 0 \rangle$ for $0 \leq t \leq 2\pi$. Consider the vector field $\mathbf{F}(x, y, z) = \langle z, x, y \rangle$, which has constant curl vector $\nabla \times \mathbf{F} = \langle 1, 1, 1 \rangle$.

- (a) Compute the tangent vectors \mathbf{r}_u and \mathbf{r}_v and their cross product $\mathbf{r}_u \times \mathbf{r}_v$.
- (b) Use part (a) to compute the flux of the vector field $\nabla \times \mathbf{F}$ across the surface D :

$$\iint_D (\nabla \times \mathbf{F}) \bullet \mathbf{N} \, dA = \iint_D (\nabla \times \mathbf{F})(\mathbf{r}(u, v)) \bullet (\mathbf{r}_u(u, v) \times \mathbf{r}_v(u, v)) \, dudv.$$

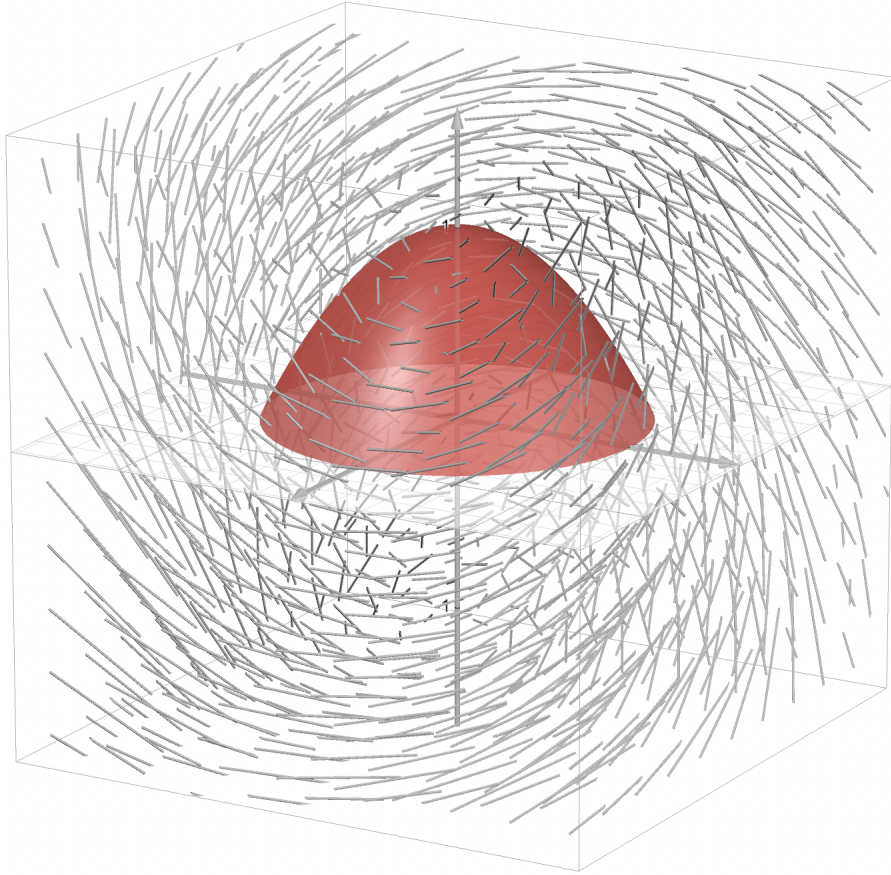
- (c) Now compute the circulation of the vector field \mathbf{F} around the boundary curve ∂D :

$$\int_{\partial D} \mathbf{F} \bullet \mathbf{T} \, ds = \int_{\partial D} \mathbf{F}(\mathbf{r}(t)) \bullet \mathbf{r}'(t) \, dt.$$

Make sure that you get the same answer as in part (a).

Here is a picture³ of the surface $\mathbf{r}(u, v)$ and the vector field \mathbf{F} . You can definitely see that this field has nonzero curl:

³You can rotate the image and play with parameters here: <https://www.desmos.com/3d/b743c9370d>



(a): First we compute the tangent vectors:

$$\begin{aligned}\mathbf{r}_u &= \langle \cos v, \sin v, -2u \rangle, \\ \mathbf{r}_v &= \langle -u \sin v, u \cos v, 0 \rangle.\end{aligned}$$

Then we compute the cross product, which gives us the “positively oriented” normal vector:

$$\begin{aligned}\mathbf{r}_u \times \mathbf{r}_v &= \langle 2u^2 \cos v, 2u^2 \sin v, u \cos^2 v + u \sin^2 v \rangle \\ &= \langle 2u^2 \cos v, 2u^2 \sin v, u \rangle.\end{aligned}$$

The unit normal vector is $\mathbf{N} = (\mathbf{r}_u \times \mathbf{r}_v) / \|\mathbf{r}_u \times \mathbf{r}_v\|$.

(b): The flux of the constant vector field $\nabla \times \mathbf{F} = \langle 1, 1, 1 \rangle$ across the surface $\mathbf{r}(u, v)$ is

$$\begin{aligned}& \iint_D (\nabla \times \mathbf{F}) \bullet \mathbf{N} \, dA \\ &= \iint_D (\nabla \times \mathbf{F})(\mathbf{r}(u, v)) \bullet \frac{\mathbf{r}_u(u, v) \times \mathbf{r}_v(u, v)}{\|\mathbf{r}_u \times \mathbf{r}_v\|} \cdot \|\mathbf{r}_u \times \mathbf{r}_v\| \, dudv \\ &= \iint_D (\nabla \times \mathbf{F})(\mathbf{r}(u, v)) \bullet (\mathbf{r}_u(u, v) \times \mathbf{r}_v(u, v)) \, dudv \\ &= \iint_D \langle 1, 1, 1 \rangle \bullet \langle 2u^2 \cos v, 2u^2 \sin v, u \rangle \, dudv \\ &= \iint_D (2u^2 \cos v + 2u^2 \sin v + u) \, dudv\end{aligned}$$

$$\begin{aligned}
&= \int_0^{2\pi} \left(\int_0^1 [2u^2 \cos v + 2u^2 \sin v + u] du \right) dv \\
&= \int_0^{2\pi} \left[2\frac{u^3}{3} \cos v + 2\frac{u^3}{3} \sin v + \frac{u^2}{2} \right]_0^1 dv \\
&= \int_0^{2\pi} \left[\frac{2}{3} \cos v + \frac{2}{3} \sin v + \frac{1}{2} \right] dv \\
&= \left[\frac{2}{3} \sin v - \frac{2}{3} \cos v + \frac{v}{2} \right]_0^{2\pi} \\
&= \left[0 - \frac{2}{3} + \frac{2\pi}{2} - 0 + \frac{2}{3} - 0 \right] \\
&= \pi.
\end{aligned}$$

(c): The circulation of $\mathbf{F} = \langle z, x, y \rangle$ around the curve $\mathbf{r}(t) = (\cos t, \sin t, 0)$ for $0 \leq t \leq 2\pi$ is

$$\begin{aligned}
\int_{\partial D} \mathbf{F} \cdot \mathbf{T} ds &= \int_0^{2\pi} \mathbf{F}(\mathbf{r}(t)) \cdot \frac{\mathbf{r}'(t)}{\|\mathbf{r}'(t)\|} \cdot \|\mathbf{r}'(t)\| dt \\
&= \int_0^{2\pi} \mathbf{F}(\mathbf{r}(t)) \cdot \mathbf{r}'(t) dt \\
&= \int_0^{2\pi} \mathbf{F}(\cos t, \sin t, 0) \cdot \langle -\sin t, \cos t, 0 \rangle dt \\
&= \int_0^{2\pi} \langle 0, \cos t, \sin t \rangle \cdot \langle -\sin t, \cos t, 0 \rangle dt \\
&= \int_0^{2\pi} (0 + \cos^2 t + 0) dt \\
&= \int_0^{2\pi} \left(\frac{1}{2} \cos(2t) + \frac{1}{2} \right) dt && \cos(2t) = 2 \cos^2 t - 1 \\
&= \left[\frac{1}{4} \cdot \sin(2t) + \frac{1}{2} \cdot t \right]_0^{2\pi} \\
&= \pi.
\end{aligned}$$

Stokes' Theorem says that the answers to (a) and (b) must be the same, and they are.