1. Various Kinds of First and Second Derivatives in $\mathbb{R}^{3}$. For any scalar field $f(x, y, z)$ we define a vector field $\operatorname{grad}(f)$ and a scalar field laplacian $(f)$ by

$$
\begin{aligned}
\operatorname{grad}(f) & =" \nabla f "=\left\langle f_{x}, f_{y}, f_{z}\right\rangle \\
\operatorname{laplacian}(f) & =" \nabla^{2} f "=f_{x x}+f_{y y}+f_{z z}
\end{aligned}
$$

and for any vector field $\mathbf{F}(x, y, z)=\langle P(x, y, z), Q(x, y, z), R(x, y, z)\rangle$ we define a vector field $\operatorname{curl}(\mathbf{F})$ and a scalar field $\operatorname{div}(\mathbf{F})$ by

$$
\begin{aligned}
\operatorname{curl}(\mathbf{F}) & =" \nabla \times \mathbf{F} "=\left\langle R_{y}-Q_{z}, P_{z}-R_{x}, Q_{x}-P_{y}\right\rangle, \\
\operatorname{div}(\mathbf{F}) & =" \nabla \bullet \mathbf{F} "=P_{x}+Q_{y}+R_{z} .
\end{aligned}
$$

(a) For any scalar field $f: \mathbb{R}^{3} \rightarrow \mathbb{R}$ check that $\operatorname{curl}(\operatorname{grad}(f))=\langle 0,0,0\rangle$.
(b) For any vector field $\mathbf{F}: \mathbb{R}^{3} \rightarrow \mathbb{R}^{3}$ check that $\operatorname{div}(\operatorname{curl}(\mathbf{F}))=0$.
(c) For any scalar field $f: \mathbb{R}^{3} \rightarrow \mathbb{R}$ check that $\operatorname{div}(\operatorname{grad}(f))=\operatorname{laplacian}(f)$.
2. Conservative Vector Fields. Consider the vector field

$$
\mathbf{F}(x, y, z)=\langle y+z, x+z, x+y\rangle .
$$

(a) Check that the curl is constantly zero: $\nabla \times \mathbf{F}(x, y, z)=\langle 0,0,0\rangle$.
(b) It follows from part (a) that there exists some scalar field $f(x, y, z)$ satisfying $\nabla f=\mathbf{F}$. Find one such scalar field. [Hint: Integrate $\mathbf{F}$ along an arbitrary path starting at some arbitrary point and ending at $(x, y, z)$.]
3. Gravitational Potential. A sun of mass $M$ sits at the origin in $\mathbb{R}^{3}$. According to Newton, the gravitational force due to the sun acting on a particle of mass $m$ at the point $(x, y, z)$ has the form

$$
\mathbf{F}(x, y, z)=\frac{-G M m}{\left(x^{2}+y^{2}+z^{2}\right)^{3 / 2}} \cdot\langle x, y, z\rangle,
$$

where $G$ is the gravitational constant.
(a) Check that the following scalar field $f: \mathbb{R}^{3} \rightarrow \mathbb{R}$ satisfies $\nabla f=\mathbf{F}$ :

$$
f(x, y, z)=\frac{+G M m}{\sqrt{x^{2}+y^{2}+z^{2}}} .
$$

(b) We can think of $U=-f$ as the gravitational potential energy. Suppose that our particle is held at rest at a position with distance $D$ from the origin. At time zero the particle is allowed to fall towards the sun. If the sun has radius $R$, use conservation of energy to compute the particle's speed when it hits the sun's surface. Your answer will involve the constants $G, M, R$ and $D$ (but not $m$ ). [Assume that $D>R$.]
4. Line Integrals and Flux Integrals in $\mathbb{R}^{2}$. Let $C$ be the parametrized path $\mathbf{r}(t)=\left(t, t^{2}\right)$ for $0 \leq t \leq 1$, with velocity vector $\mathbf{r}^{\prime}(t)=\langle 1,2 t\rangle$. From this parametrization we can define a unit tangent vector and a unit normal vector to the curve $C$ at the point $\mathbf{r}(t)$ :

$$
\mathbf{T}=\frac{\langle 1,2 t\rangle}{\sqrt{1+4 t^{2}}} \quad \text { and } \quad \mathbf{N}=\frac{\langle 2 t,-1\rangle}{\sqrt{1+4 t^{2}}}
$$

Now consider the constant vector field $\mathbf{F}(x, y)=\langle 3,1\rangle$.
(a) Compute the line integral $\int_{C} \mathbf{F} \bullet \mathbf{T} d s$.
(b) Compute the flux integral $\int_{C} \mathbf{F} \bullet \mathbf{N} d s$.
5. Green's Theorem on a Circle. Let $D$ be the unit disk in $\mathbb{R}^{2}$ centered at $(0,0)$. Consider the vector field $\mathbf{F}=\langle P, Q\rangle=\left\langle x y^{2}, x+y\right\rangle$ with $\operatorname{curl}(\mathbf{F})=Q_{x}-P_{y}=1-2 x y$.
(a) Compute the integral $\iint_{D} \operatorname{curl}(\mathbf{F}) d A$. [Hint: Polar coordinates are easiest. You may use the trigonometric identity $\sin (2 \theta)=2 \sin \theta \cos \theta$.]
(b) The boundary curve $\partial D$ is the unit circle, oriented counterclockwise. Use the standard parametrization $\mathbf{r}(t)=(\cos t, \sin t)$ with $0 \leq t \leq 2 \pi$ to set up the integral $\oint_{\partial D} \mathbf{F} \bullet \mathbf{T} d s$. You will probably not be able to evaluate this integral by hand. Use a computer to verify that you get the same answer as in part (a).
6. Stokes' Theorem on a Parabolic Dome. Let $D$ be the two-dimensional surface in $\mathbb{R}^{3}$ defined by $z=1-x^{2}-y^{2}$ and $z \geq 0$. This surface can be parametrized by

$$
\mathbf{r}(u, v)=\left\langle u \cos v, u \sin v, 1-u^{2}\right\rangle \quad \text { with } 0 \leq u \leq 1 \text { and } 0 \leq v \leq 2 \pi .
$$

The boundary curve $\partial D$ is the unit circle in the $x, y$-plane, oriented counterclockwise, which can be parametrized as $\mathbf{r}(t)=\langle\cos t, \sin t, 0\rangle$ for $0 \leq t \leq 2 \pi$. Consider the vector field $\mathbf{F}(x, y, z)=\langle z, x, y\rangle$, which has constant curl vector $\nabla \times \mathbf{F}=\langle 1,1,1\rangle$.
(a) Compute the tangent vectors $\mathbf{r}_{u}$ and $\mathbf{r}_{v}$ and their cross product $\mathbf{r}_{u} \times \mathbf{r}_{v}$.
(b) Use part (a) to compute the flux of the vector field $\nabla \times \mathbf{F}$ across the surface $D$ :

$$
\iint_{D}(\nabla \times \mathbf{F}) \bullet \mathbf{N} d A=\iint_{D}(\nabla \times \mathbf{F})(\mathbf{r}(u, v)) \bullet\left(\mathbf{r}_{u}(u, v) \times \mathbf{r}_{v}(u, v)\right) d u d v
$$

(c) Now compute the circulation of the vector field $\mathbf{F}$ around the boundary curve $\partial D$ :

$$
\int_{\partial D} \mathbf{F} \bullet \mathbf{T} d s=\int_{\partial D} \mathbf{F}(\mathbf{r}(t)) \bullet \mathbf{r}^{\prime}(t) d t
$$

Make sure that you get the same answer as in part (a).

