# Vacuum Spacetimes with Compact Cauchy Horizons with Non-Closed Generators

JIM ISENBERG (UNIVERSITY OF OREGON) VINCE MONCRIEF (YALE UNIVERSITY)

# Taub-NUT Spacetimes

- Misner: Taub-NUT Spacetimes as the counter-example to almost everything
- Globally hyperbolic region extends across Cauchy horizon
- Cauchy horizon indicates breakdown of determinism
- Taub-NUT Cauchy horizons are compact

# Diagram of Taub-NUT Space-Time



# Strong Cosmic Censorship Conjecture

- Generic solutions of Einstein's equations do not contain Cauchy horizons
- There are an infinite number of space-times which do have Cauchy horizons
- ? Do Cauchy horizons occur in generic solutions of Einstein's equations?
- Po compact Cauchy horizons occur in generic solutions of Einstein's equations?
- This is the focus of my talk here

# THEOREM (Isenberg and Moncrief)

- Space-time solutions of Einstein's Equation with compact Cauchy horizons with non-degenerate generators must contain a Killing vector field
- Previous results required generators of the horizon to be closed.
- Our new results allow the generators to be non-closed (but not ergodic).
- Also, if the generators are not closed, there must be a twodimensional isometry group: i.e. two independent Killing vector fields

# Key Ingredients of the Proof

- I) Topology of the Compact Cauchy Horizon and its Non-Closed Generators
- 2) Adapted Geodesic Null Coordinates
- 3) Analyticity
- 4) Einstein Equations
- 5) Poincaré Recurrence

# Key Ingredients of the Proof (continued)

- 6) Invariance of Geometry of Discs Transverse to Generators
- 7) ''Ribbon Arguments'' using Stokes'Theorem
- 8) Kolmogorov–Kontsevich Theorem on Transverse Foliation of a 2-Torus
- 9) Grauert Tubes and the Banach Space of Complex Analytic Functions
- 10) Cauchy-Kowaleski Theorem

#### Basic Steps of the Proof

• I) Topology of the Cauchy Horizon and its Non-Closed Generators







 2) Choose Adapted Geodesic Null Coordinates for Null Surface N Corresponding to Zero Level Set of Analytic Function t and Null Vector Field K

$$g = dt \otimes dx^{3} + dx^{3} \otimes dt$$
$$+ \varphi \, dx^{3} \otimes dx^{3} + \beta_{a}(dx^{a} \otimes dx^{3} + dx^{3} \otimes dx^{a})$$
$$+ \mu_{ab} \, dx^{a} \otimes dx^{b}.$$

$$\varphi|_{t=0} = \beta_a|_{t=0} = 0$$

• 3) Use Hawking & Ellis Result Showing that the Volume Element of the Transverse Discs is Invariant Along the Generator Flow.

$$\left(\det \mu_{ab}\right)_{,3}\big|_{t=0} = 0$$

• 4) Use Einstein's equations to Show that the Metric on the Transverse Discs is Invariant along the Generator Flow

$$R_{tt} = -\frac{1}{2} \mu^{ab} \mu_{ab,tt} + \frac{1}{4} \mu^{ac} \mu^{bd} \mu_{ab,t} \mu_{cd,t},$$

$$\begin{split} R_{t3} &= \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \bigg( \frac{1}{2} \phi_{,t} - \frac{1}{2} \beta^a \beta_{a,t} \bigg) \bigg]_{,t} + \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \bigg( \frac{\mu^{ab}}{2} \beta_{a,t} \bigg) \bigg]_{,b} \\ &- \frac{1}{2} \mu^{ab} \mu_{ab,t3} + \frac{1}{4} \mu^{ac} \mu^{bd} \mu_{ab,t} \mu_{cd,3} \,, \end{split}$$

$$R_{tb} = \frac{1}{\sqrt{\mu}} \left[ \sqrt{\mu} \left( \frac{1}{2} \beta_{b,t} - \frac{1}{2} \beta^{a} \mu_{ab,t} \right) \right]_{,t} + {}^{(2)} \nabla_{c} \left( \frac{1}{2} \mu^{ac} \mu_{ab,t} \right) - {}^{(2)} \nabla_{b} \left( \frac{1}{2} \mu^{cd} \mu_{cd,t} \right),$$

$$\begin{split} R_{ab} &= \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \bigg( -\frac{1}{2} \mu_{ab,3} + \frac{1}{2} ({}^{(2)}\nabla_{b}\beta_{a} + {}^{(2)}\nabla_{a}\beta_{b}) + \frac{\phi}{2} \mu_{ab,t} - \frac{1}{2} \beta^{c}\beta_{c}\mu_{ab,t} \bigg) \bigg]_{,t} \\ &+ {}^{(2)}\nabla_{c} \bigg[ \frac{\beta^{c}}{2} \mu_{ab,t} \bigg] - \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \frac{\mu_{ab,t}}{2} \bigg]_{,3} \\ &+ {}^{(2)}R_{ab} - \{ \frac{1}{2} (\beta_{a,t} - \beta^{c}\mu_{ac,t}) (\beta_{b,t} - \beta^{d}\mu_{bd,t}) \\ &+ \frac{1}{4} \mu^{df} \mu_{fb,t} (-2\mu_{ad,3} + 2 {}^{(2)}\nabla_{d}\beta_{a} + (\phi - \beta^{g}\beta_{g}) \mu_{ad,t}) \\ &+ \frac{1}{4} \mu^{df} \mu_{fa,t} (-2\mu_{bd,3} + 2 {}^{(2)}\nabla_{d}\beta_{b} + (\phi - \beta^{g}\beta_{g}) \mu_{bd,t}) \} \,, \end{split}$$

$$\begin{split} R_{3b} &= \frac{1}{\sqrt{\mu}} \left[ \sqrt{\mu} \left( \frac{1}{2} \phi_{,b} + \frac{1}{2} \phi_{\beta_{b,t}} - \frac{1}{2} \beta^{a} (\beta_{a,b} - \beta_{b,a}) - \frac{1}{2} \beta^{a} \beta_{a} \beta_{b,t} - \frac{1}{2} \beta^{a} \mu_{ab,3} \right) \right]_{,t} \\ &+ {}^{(2)} \nabla_{c} \left[ \frac{\mu^{ac}}{2} (\beta_{a,b} - \beta_{b,a}) + \frac{1}{2} \beta^{c} \beta_{b,t} + \frac{1}{2} \mu^{ac} \mu_{ab,3} \right] \\ &- \frac{1}{\sqrt{\mu}} \left[ \sqrt{\mu} \frac{\beta_{b,t}}{2} \right]_{,3} - \frac{1}{2} \mu^{cd} {}^{(2)} \nabla_{b} (\mu_{cd,3}) \\ &- \left\{ \frac{1}{2} (\beta_{b,t} - \beta^{a} \mu_{ab,t}) (\phi_{,t} - \beta^{c} \beta_{c,t}) \right. \\ &+ \frac{1}{2} \mu^{dc} \mu_{cb,t} (\phi_{,d} + \phi \beta_{d,t} - \beta^{a} (\beta_{a,d} - \beta_{d,a}) - \beta_{d,3} - \beta^{a} \beta_{a} \beta_{d,t}) \\ &+ \frac{1}{2} \mu^{ac} \beta_{a,t} (-\mu_{bc,3} + {}^{(2)} \nabla_{c} \beta_{b}) \right\}, \end{split}$$

$$\begin{split} R_{33} &= \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \bigg( \frac{1}{2} \phi_{,3} + \frac{1}{2} \phi \phi_{,t} + \frac{1}{2} \beta^{a} \phi_{,a} - \beta^{a} \beta_{a,3} - \frac{1}{2} \beta^{a} \beta_{a} \phi_{,t} \bigg) \bigg]_{,t} \\ &+ \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \, \mu^{ac} \bigg( -\frac{\phi_{,a}}{2} + \beta_{a,3} + \frac{\beta_{a}}{2} \phi_{,t} \bigg) \bigg]_{,c} \\ &- \frac{1}{\sqrt{\mu}} \bigg[ \sqrt{\mu} \frac{\phi_{,t}}{2} \bigg]_{,3} - \frac{1}{2} \mu^{ab} \mu_{ab,33} + \frac{1}{4} \mu^{ac} \mu^{bd} \mu_{ab,3} \mu_{cd,3} \\ &- \{ \frac{1}{2} (\phi_{,t} - \beta^{a} \beta_{a,t})^{2} + \frac{1}{4} \mu^{ac} \mu^{bd} (\beta_{a,b} - \beta_{b,a}) (\beta_{d,c} - \beta_{c,d}) \\ &+ \frac{1}{2} \mu^{dc} \beta_{c,t} (2\phi_{,d} + \phi \beta_{d,t} - 2\beta^{a} (\beta_{a,d} - \beta_{d,a}) - 2\beta_{d,3} - \beta^{a} \beta_{a} \beta_{d,t} ) \big\} \,, \end{split}$$

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# R33 = 0, Restricted to N

$$\begin{split} \mathring{R}_{33} &= 0 = \left[ \left( \ell n \sqrt{\det \mu} \right)_{,33} \\ &+ \frac{1}{2} \varphi_{,t} \left( \ell n \sqrt{\det \mu} \right)_{,3} \\ &+ \frac{1}{4} \mu^{ac} \mu^{bd} \mu_{ab,3} \mu_{cd,3} \right] \Big|_{t=0} \end{split}$$

• 5) Construction of Fiducial Riemannian Metric on the Space-Time

$$g'(Y,Z) = g(Y,Z) + 2g(Y,V)g(Z,V)$$

• 6) Restrictions on the Geometry of the Transverse Discs following from Poincaré Recurrence and Invariance of Geometry along the Flow

• Scalar curvature is constant on the horizon N.

• 7) Closures of the Orbits of the Geodesics Form 2-Tori

• 8) Kolmogorov-Kontsevich Theorem Guarantees Existence of Foliations of Geodesic 2-Tori by Transverse Circles

- 9) Define Privileged One-form  $\boldsymbol{\omega}$  Based on Space-Time Connection
- Expressions for  $\boldsymbol{\omega}$  and its Exterior Derivative can be Expressed in Terms of the Metric Coefficients and their Derivatives:

$$\omega_X = -\frac{1}{2}\mathring{\varphi}_{,t}dx^3 - \frac{1}{2}\mathring{\beta}_{a,t}dx^a.$$

$$d\omega_X = -\frac{1}{2}(\mathring{\varphi}_{,ta} - \mathring{\beta}_{a,t3})dx^a \wedge dx^3 - \frac{1}{2}\mathring{\beta}_{a,tb}dx^b \wedge dx^a.$$

• 10) Use Einstein's Equations to Simplify Formula for  $d\omega$ :

$$d\omega_X = -\frac{1}{2}\mathring{\beta}_{a,tb}dx^b \wedge dx^a.$$

• 11) Construct Ribbon from Path on N Flowed along Geodesics through the Ends of the Path



• 12) Use Stokes'Theorem to Show that the Integral over the Boundary of the Ribbon of the One-form  $\boldsymbol{\omega}$  vanishes.

$$\int_{\partial r} \omega_X = \int_r d\omega_X = 0$$

- 13) Use Ribbon Argument to Correlate Integrals of  $\boldsymbol{\omega}$  along Neighboring Generators

- 14) Use Ribbon Argument to Show that all Generators are either Complete or Incomplete:
- We restrict to the Case that the Generators are All Complete in One Direction and Incomplete in the Other Direction
- We call this the Non-Degenerate Case

- I5) Candidate Vector Field Defined as Function U times Vector Field Tangent to Generators
- The formula is given by

$$u(x^3, x^a) = \frac{k}{2} \int_{x^3}^{\infty} d\rho \exp\left[-\int_{x^3}^{\rho} \frac{\mathring{\varphi}_{,t}}{2}(\xi, x^a)d\xi\right].$$

• 16) A Cauchy Sequence for U:

$$u_i(x^3, x^a) = \frac{k}{2} \int_{x^3}^{x^3 + is^*} d\rho \exp\left[-\int_{x^3}^{\rho} \frac{\mathring{\varphi}_{,t}}{2}(\xi, x^a)d\xi\right]$$

- 17) Complexify the Space Using Grauert Tubes and Show that the Cauchy Sequence Converges:
- The Limit is an Analytic Function U

 I8) Initial Data for Cauchy-Kowaleski on the Cauchy Horizon N for Expansion into the Space-Time of the Killing Vector Field.

$$\mu_{a'b',t'x^{3'}}\bigg|_{t'=0} = \left.\left(\frac{\partial x^c}{\partial x^{a'}}\frac{\partial x^d}{\partial x^{b'}}\mu_{cd,t3}\right)\bigg|_{t=0}$$

• 19) Patch Together Space-Time Geometry with Killing Vector Field

 20) Use Isenberg-Moncrief Results to Show that a Killing Field with Non-Closed Orbits on a Compact Manifold Must Generate a Two-Dimensional Isometry Group

#### Extensions of this work?

- What about Non-Analytic Space-Times?
  - We don't know
- What about Degenerate Cauchy Horizons?
  - We don't think Space-Times with Degenerate Cauchy Horizons Exist
- What about Cauchy Horizons with Ergodic Generators?
  - We think that these are all Kasner Space-Times with Nonrational Identifications